Symmetries of the gauge theories in the light front

Olivera Mišković

Pontificia Universidad Católica de Valparaíso, Chile

Collaborators Oriana Labrin (PUCV, Chile), Hernán González (UAI, Chile)

arXiv: 2302.xxxx



Workshop on Gravity, Holography, Strings & Noncommutative Geometry 3rd February 2023, Belgrade, Serbia

1 Infrared structure of gauge theories

E

イロト イヨト イヨト イヨト

- 1 Infrared structure of gauge theories
- **2** Hamiltonian analysis of electromagnetism in the null foliation

3

・ロト ・回 ト ・ ヨト ・

- 1 Infrared structure of gauge theories
- 2 Hamiltonian analysis of electromagnetism in the null foliation
- **3** New symmetry generator; Beyond U(1)

- 1 Infrared structure of gauge theories
- 2 Hamiltonian analysis of electromagnetism in the null foliation
- **3** New symmetry generator; Beyond U(1)
- Extension to Yang-Mills theory

- 1 Infrared structure of gauge theories
- 2 Hamiltonian analysis of electromagnetism in the null foliation
- **3** New symmetry generator; Beyond U(1)
- Extension to Yang-Mills theory
- **5** Discussion

IR region of theories with massless particles in asymptotically flat spaces

IR region of theories with massless particles in asymptotically flat spaces



Motivation

- Hamiltonian treatment of asymptotic symmetries [Bondi, van der Burg, Metzner 1962; Sachs 1962]
- **BMS symmetry** infinite-dimensional asymptotic symmetry at the null boundary of 4D asymptotically flat spacetimes

• Celestial holography

- Duality between massless 4D asymptotically flat spacetimes and 2D CFT on the celestial sphere

Motivation

- Hamiltonian treatment of asymptotic symmetries [Bondi, van der Burg, Metzner 1962; Sachs 1962]
- **BMS symmetry** infinite-dimensional asymptotic symmetry at the null boundary of 4D asymptotically flat spacetimes

• Celestial holography

- Duality between massless 4D asymptotically flat spacetimes and 2D CFT on the celestial sphere
- Asymptotic symmetries in Electromagnetic and Yang-Mills theories
- 2D realization of soft symmetries in electromagnetism

[He, Mitra, Porfyriadis, Strominger 2014; Nande, Pate, Strominger 2018]

• Extension to Yang-Mills [Strominger 2014; He, Mitra, Strominger, 2016]

Э

イロト イヨト イヨト イヨト

Vacuum degeneracy in gauge theories $(\omega \rightarrow 0)$

 \Leftrightarrow Enhancement of symmetries at the boundary of flat spacetime $(r \to \infty)$

Vacuum degeneracy in gauge theories $(\omega \rightarrow 0)$

 \Leftrightarrow Enhancement of symmetries at the boundary of flat spacetime $(r
ightarrow \infty)$

• Goldstone modes, do not leave the vacuum invariant: $e^{S[\eta]} |\psi\rangle = |\psi + \eta\rangle$

Infrared structure of gauge theories

Vacuum degeneracy in gauge theories $(\omega \rightarrow 0)$

 \Leftrightarrow Enhancement of symmetries at the boundary of flat spacetime $(r
ightarrow \infty)$

• Goldstone modes, do not leave the vacuum invariant: $e^{S[\eta]} |\psi\rangle = |\psi + \eta\rangle$



5 / 30

Infrared structure of gauge theories

Penrose diagram of the Minkowski space



6 / 30

Null foliated reference frame

• Minkowski metric in D = 4 in the spherical coordinates (t, r, y^A)

 $M_4: ds^2 = -dt^2 + dr^2 + r^2 d\Omega^2$ $S^2: d\Omega^2 = \gamma_{AB}(y) dy^A dy^B$

Null foliated reference frame

• Minkowski metric in D = 4 in the spherical coordinates (t, r, y^A)

 $M_4: ds^2 = -dt^2 + dr^2 + r^2 d\Omega^2$

 $\mathbb{S}^2: \quad \mathrm{d}\Omega^2 = \gamma_{AB}(y)\,\mathrm{d}y^A\mathrm{d}y^B$

- Time coordinate $u = t \epsilon r$, $-1 \le \epsilon \le 1$
 - $\epsilon = 1$ retarded time
 - $\epsilon = 0$ proper time of a massive particle
 - $\epsilon = -1$ advanced time

3

7 / 30

Null foliated reference frame

• Minkowski metric in D = 4 in the spherical coordinates (t, r, y^A)

 $M_4: \quad \mathrm{d}s^2 = -\mathrm{d}t^2 + \mathrm{d}r^2 + r^2\mathrm{d}\Omega^2$

 \mathbb{S}^2 : $\mathrm{d}\Omega^2 = \gamma_{AB}(y) \,\mathrm{d}y^A \mathrm{d}y^B$

- Time coordinate $u = t \epsilon r$, $-1 \le \epsilon \le 1$
 - $\epsilon = 1$ retarded time
 - $\epsilon = 0$ proper time of a massive particle
 - $\epsilon = -1$ advanced time
- Coordinates on S²: stereographic projection $(\theta, \varphi) \rightarrow y^{A} = (z, \bar{z})$

$$z={
m e}^{{
m i}arphi}\cotrac{ heta}{2}$$
 , $ar z={
m e}^{-{
m i}arphi}\cotrac{ heta}{2}$



• Minkowski metric $g_{\mu\nu}$ in the coordinates $x^{\mu} = (u, r, y^{A})$:

$$\mathrm{d} s^2 = -\mathrm{d} u^2 - 2\epsilon \,\mathrm{d} u \mathrm{d} r + \left(1 - \epsilon^2\right) \mathrm{d} r^2 + r^2 \mathrm{d} \Omega^2$$

• \mathbb{S}^2 metric in complex coordinates

$$\gamma_{AB} = \left(egin{array}{cc} 0 & \gamma_{zar{z}} \ \gamma_{zar{z}} & 0 \end{array}
ight), \quad \gamma_{zar{z}} = rac{2}{(1+zar{z})^2} = \sqrt{\gamma}$$

8 / 30

Image: A match a ma

- Electromagnetic action in the background $\mathfrak{g}_{\mu\nu}$

$$I[A] = -\frac{1}{4e^2} \int d^4 x \sqrt{\mathfrak{g}} F^{\mu\nu} F_{\mu\nu} \qquad (F_{\mu\nu} = \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu})$$

• Canonical momenta $\pi^{\mu} = -\frac{1}{e^2} \sqrt{\mathfrak{g}} F^{\mu\mu}$

9 / 30

<ロ> <四> <四> <三> <三> <三> <三>

• Electromagnetic action in the background $\mathfrak{g}_{\mu\nu}$

$$I[A] = -\frac{1}{4e^2} \int d^4 x \sqrt{\mathfrak{g}} F^{\mu\nu} F_{\mu\nu} \qquad (F_{\mu\nu} = \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu})$$

• Canonical momenta $\pi^{\mu} = -\frac{1}{e^2} \sqrt{\mathfrak{g}} F^{\mu\mu}$

In components:

$$\begin{aligned} \pi^{u} &= 0 \\ \pi^{r} &= \frac{r^{2}}{e^{2}} \sqrt{\gamma} F_{ur} \\ \pi^{A} &= -\frac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} \left[(\epsilon^{2} - 1) F_{uB} - \epsilon F_{rB} \right] \end{aligned}$$

9 / 30

• Electromagnetic action in the background $\mathfrak{g}_{\mu\nu}$

$$I[A] = -\frac{1}{4e^2} \int d^4 x \sqrt{\mathfrak{g}} F^{\mu\nu} F_{\mu\nu} \qquad (F_{\mu\nu} = \partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu})$$

• Canonical momenta $\pi^{\mu} = -\frac{1}{e^2} \sqrt{\mathfrak{g}} F^{\mu\mu}$

In components:

$$\begin{aligned} \pi^{u} &= 0 \\ \pi^{r} &= \frac{r^{2}}{e^{2}} \sqrt{\gamma} F_{ur} \\ \pi^{A} &= -\frac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} \left[(\epsilon^{2} - 1) F_{uB} - \epsilon F_{rB} \right] \end{aligned}$$

- The limit $\epsilon^2
 ightarrow 1$ is discontinuous
- The action in the light-cone $(\epsilon^2 = 1)$ has an additional constraint
- Light-cone actions are first order in velocities [Steinhardt 1980]

9 / 30

Bondi reference frame: $\epsilon^2 = 1$

• Primary constraints

$$\pi^{u} pprox 0$$
, $\chi^{A} \equiv \epsilon \pi^{A} - rac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} F_{rB} pprox 0$

Image: A match a ma

Bondi reference frame: $\epsilon^2 = 1$

• Primary constraints

$$\pi^{u} pprox 0$$
 , $\chi^{A} \equiv \epsilon \pi^{A} - rac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} F_{rB} pprox 0$

- Canonical Hamiltonian $\mathcal{H}_{C}=\pi^{\mu}\dot{A}_{\mu}-\mathcal{L}$

$$\mathcal{H}_{C} = rac{e^{2}(\pi^{r})^{2}}{2r^{2}\sqrt{\gamma}} + rac{e^{2} ilde{\pi}_{A}\pi^{A}}{2\sqrt{\gamma}} + rac{\sqrt{\gamma}}{4e^{2}r^{2}} ilde{F}^{AB}F_{AB} - A_{u}\partial_{i}\pi^{i}$$

Bondi reference frame: $\epsilon^2 = 1$

• Primary constraints

$$\pi^{u} pprox 0$$
 , $\chi^{A} \equiv \epsilon \pi^{A} - rac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} F_{rB} pprox 0$

• Canonical Hamiltonian $\mathcal{H}_{C}=\pi^{\mu}\dot{A}_{\mu}-\mathcal{L}$

$$\mathcal{H}_{\mathcal{C}} = rac{e^2(\pi^r)^2}{2r^2\sqrt{\gamma}} + rac{e^2 ilde{\pi}_A\pi^A}{2\sqrt{\gamma}} + rac{\sqrt{\gamma}}{4e^2r^2}\, ilde{\mathcal{F}}^{AB}\mathcal{F}_{AB} - \mathcal{A}_u\partial_i\pi^i$$

• Total Hamiltonian [Dirac 1964]

 $\mathcal{H}_{T} = \mathcal{H}_{C} + \lambda_{u} \pi^{u} + \lambda_{A} \chi^{A}$

 $\lambda_u(x), \, \lambda_A(x) =$ Hamiltonian multipliers

Bondi reference frame: $\epsilon^2 = 1$

• Primary constraints

$$\pi^{u} \approx 0, \quad \chi^{A} \equiv \epsilon \pi^{A} - \frac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} F_{rB} \approx 0$$

• Canonical Hamiltonian $\mathcal{H}_{C}=\pi^{\mu}\dot{A}_{\mu}-\mathcal{L}$

$$\mathcal{H}_{\mathcal{C}} = rac{e^2(\pi^r)^2}{2r^2\sqrt{\gamma}} + rac{e^2 ilde{\pi}_A\pi^A}{2\sqrt{\gamma}} + rac{\sqrt{\gamma}}{4e^2r^2}\, ilde{\mathcal{F}}^{AB}\mathcal{F}_{AB} - \mathcal{A}_u\partial_i\pi^i$$

• Total Hamiltonian [Dirac 1964]

 $\mathcal{H}_{T} = \mathcal{H}_{C} + \lambda_{u}\pi^{u} + \lambda_{A}\chi^{A}$

 $\lambda_u(x)$, $\lambda_A(x) =$ Hamiltonian multipliers

• Canonical Poisson brackets $\{A_{\mu}(x), \pi^{\nu}(x')\}_{\mu=\mu'} = \delta^{\nu}_{\mu} \delta^{(3)}(x-x')$

• Evolution $\dot{\Phi}(x) \approx \{\Phi(x), H_T\} = \int d^3x' \{\Phi(x), \mathcal{H}_T(x')\}$

< □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > <

Comment

- We always have two different Hamiltonians $\epsilon = \pm 1$.
- The future cone $\epsilon = +1$ and the past cone $\epsilon = -1$ are causally disconnected.
- Common boundary i^0 is located at the spatial infinity.

Comment

- We always have two different Hamiltonians $\epsilon = \pm 1$.
- The future cone $\epsilon = +1$ and the past cone $\epsilon = -1$ are causally disconnected.
- Common boundary *i*⁰ is located at the spatial infinity.
- A radial wave located at (u, r) propagates along $v = t + \epsilon r = const$.
- In absence of massive particles, the future lightlike infinity $\mathcal{J}^+ = \mathcal{J}^+(t-r,y)$ and the past lightlike infinity $\mathcal{J}^- = \mathcal{J}^-(t+r,y)$ behave as Cauchy surfaces [Hawking, Perry, Strominger 2016].



< □ > < 同 >

Symplectic matrix

$$\left\{\chi^{A}(x),\chi^{B}(x')\right\} = \Omega^{AB}(x,x') \equiv -\frac{2\epsilon}{e^{2}}\sqrt{\gamma}\gamma^{AB}\partial_{r}\delta^{(3)}$$

Image: A match a ma

Symplectic matrix

$$\left\{\chi^{A}(x),\chi^{B}(x')\right\} = \Omega^{AB}(x,x') \equiv -\frac{2\epsilon}{e^{2}}\sqrt{\gamma}\gamma^{AB}\partial_{r}\delta^{(3)}$$

- If Ω^{AB} invertible: χ^A are **first class** (generate symmetries)
- If Ω^{AB} not invertible: χ^A are second class (eliminate redundant fields)

Symplectic matrix

$$\left\{\chi^{A}(x),\chi^{B}(x')\right\} = \Omega^{AB}(x,x') \equiv -\frac{2\epsilon}{e^{2}}\sqrt{\gamma}\gamma^{AB}\partial_{r}\delta^{(3)}$$

- If Ω^{AB} invertible: χ^A are **first class** (generate symmetries)
- If Ω^{AB} not invertible: χ^A are second class (eliminate redundant fields)

One possibility

- Ω^{AB} is invertible because γ^{AB} is invertible $\Rightarrow \chi^A$ are second class
- Reduced phase space $\chi^A = 0$ [Goldberg 1991, Majumdar 2022]

Symplectic matrix

$$\left\{\chi^{A}(x),\chi^{B}(x')\right\} = \Omega^{AB}(x,x') \equiv -\frac{2\epsilon}{e^{2}}\sqrt{\gamma}\gamma^{AB}\partial_{r}\delta^{(3)}$$

- If Ω^{AB} invertible: χ^A are **first class** (generate symmetries)
- If Ω^{AB} not invertible: χ^A are second class (eliminate redundant fields)

One possibility

- Ω^{AB} is invertible because γ^{AB} is invertible $\Rightarrow \chi^A$ are second class
- Reduced phase space $\chi^A = 0$ [Goldberg 1991, Majumdar 2022]

Other possibility

- Ω^{AB} is invertible, but its inverse is not unique
- Ω^{AB} is infinite-dimensional matrix and it has zero modes

 $\int \mathrm{d}^3 x' \,\Omega^{AB} V'_B = -\tfrac{2\epsilon}{e^2} \sqrt{\gamma} \gamma^{AB} \partial_r V_B = 0 \quad \Rightarrow \quad \bigvee_{A \equiv b} V_B = V_B(y)$

Symplectic matrix

$$\left\{\chi^{A}(x),\chi^{B}(x')\right\} = \Omega^{AB}(x,x') \equiv -\frac{2\epsilon}{e^{2}}\sqrt{\gamma}\gamma^{AB}\partial_{r}\delta^{(3)}$$

- If Ω^{AB} invertible: χ^A are **first class** (generate symmetries)
- If Ω^{AB} not invertible: χ^A are second class (eliminate redundant fields)

One possibility

- Ω^{AB} is invertible because γ^{AB} is invertible $\Rightarrow \chi^A$ are second class
- Reduced phase space $\chi^{A} = 0$ [Goldberg 1991, Majumdar 2022]

Other possibility

- Ω^{AB} is invertible, but its inverse is not unique
- Ω^{AB} is infinite-dimensional matrix and it has zero modes

 $\Rightarrow \chi^{A}_{(0)}(y)$ is first class constraint (*r*-independent part of the constraint)

13 / 30

Consistency conditions

(日) (同) (三) (

Consistency conditions

• Conservation of constraints during their evolution

$$\begin{split} \dot{\pi}^{u} &= 0 \qquad \Rightarrow \quad \chi = \partial_{i}\pi^{i} \approx 0 \quad \text{(differential Gauss law)} \\ \dot{\chi}^{A} &= 0 \qquad \Rightarrow \quad \int d^{3}x' \,\Omega^{AB}\lambda'_{B} = \int d^{3}x' \,\{\mathcal{H}'_{C},\chi^{A}\} \end{split}$$

Consistency conditions

• Conservation of constraints during their evolution

$$\begin{split} \dot{\pi}^{u} &= 0 \qquad \Rightarrow \quad \chi = \partial_{i} \pi^{i} \approx 0 \quad \text{(differential Gauss law)} \\ \dot{\chi}^{A} &= 0 \qquad \Rightarrow \quad \int d^{3} x' \, \Omega^{AB} \lambda'_{B} = \int d^{3} x' \left\{ \mathcal{H}'_{C}, \chi^{A} \right\} \end{split}$$

• The multiplier λ_A is not fully determined

$$\partial_r \lambda_A = -\frac{\epsilon e^2}{2\sqrt{\gamma}} \partial_r \tilde{\pi}_A - \frac{1}{2r^2} \nabla^B F_{AB} + \frac{\epsilon e^2}{2r^2} \partial_B \left(\frac{\pi^r}{\sqrt{\gamma}}\right)$$
$$\lambda_A = \bar{\lambda}_A + \Lambda_A(y)$$

- A free function $\Lambda_A(y)$ is due to the zero modes of Ω^{AB}
- $\bar{\lambda}_A$ determined part of λ_A

イロト イポト イヨト イヨト 二日

Summary of the constraints

 $\begin{array}{ll} \text{Primary constraints:} & \pi^{u} \,, \quad \chi^{A} = \epsilon \pi^{A} - \frac{1}{e^{2}} \sqrt{\gamma} \gamma^{AB} F_{rB} \\ \text{Secondary constraint:} & \chi = \partial_{i} \pi^{i} \,. \end{array}$
Summary of the constraints

 $\begin{array}{ll} \mbox{Primary constraints:} & \pi^u \,, & \chi^A = \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB} \\ \mbox{Secondary constraint:} & \chi = \partial_i \pi^i \,. \end{array}$

Class of the constraints

Summary of the constraints

 $\begin{array}{ll} \mbox{Primary constraints:} & \pi^u \,, & \chi^A = \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB} \\ \mbox{Secondary constraint:} & \chi = \partial_i \pi^i \,. \end{array}$

Class of the constraints

• π^u – first class, A_u behaves as a multiplier in the Hamiltonian

Summary of the constraints

 $\begin{array}{ll} \mbox{Primary constraints:} & \pi^u \,, & \chi^A = \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB} \\ \mbox{Secondary constraint:} & \chi = \partial_i \pi^i \,. \end{array}$

Class of the constraints

- π^u first class, A_u behaves as a multiplier in the Hamiltonian
- χ first class, differential Gauss law, $\pi^i = \sqrt{\gamma} E^i$

イロト イポト イヨト イヨト

Summary of the constraints

 $\begin{array}{ll} \mbox{Primary constraints:} & \pi^u \,, & \chi^A = \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB} \\ \mbox{Secondary constraint:} & \chi = \partial_i \pi^i \,. \end{array}$

Class of the constraints

- π^u first class, A_u behaves as a multiplier in the Hamiltonian
- χ first class, differential Gauss law, $\pi^i = \sqrt{\gamma} E^i$
- $\chi^{A}_{(0)}$ first class, r-independent part of the constraint

Summary of the constraints

 $\begin{array}{ll} \mbox{Primary constraints:} & \pi^u \,, & \chi^A = \epsilon \pi^A - \frac{1}{e^2} \sqrt{\gamma} \gamma^{AB} F_{rB} \\ \mbox{Secondary constraint:} & \chi = \partial_i \pi^i \,. \end{array}$

Class of the constraints

- π^u first class, A_u behaves as a multiplier in the Hamiltonian
- χ first class, differential Gauss law, $\pi^i = \sqrt{\gamma} E^i$
- $\chi^{A}_{(0)}$ first class, r-independent part of the constraint
- $\chi^{\mathcal{A}}_{(n)}$ $(n\geq 1)$ second class, coefficients of the Taylor expansion in $rac{1}{r}$
- We have to expand all the fields asymptotically in the vicinity of the boundary r = const → ∞.

Standard asymptotic conditions of the fields [Strominger 2014]

$$\begin{aligned} A_u &= \mathcal{O}(\frac{1}{r}), \qquad A_r &= \mathcal{O}(\frac{1}{r^2}), \qquad A_A &= \mathcal{O}(r^0), \\ \pi^u &= 0, \qquad \pi^r &= \mathcal{O}(r^0), \qquad \pi^A &= \mathcal{O}(\frac{1}{r^2}) \end{aligned}$$

• Boundary fields: $A_{(0)A} = A_A|_{r \to \infty}$, $\pi^r_{(0)} = \pi^r|_{r \to \infty}$

イロト イポト イヨト イ

(PUCV)

16 / 30

Standard asymptotic conditions of the fields [Strominger 2014]

$$\begin{aligned} A_u &= \mathcal{O}(\frac{1}{r}), \qquad A_r &= \mathcal{O}(\frac{1}{r^2}), \qquad A_A &= \mathcal{O}(r^0), \\ \pi^u &= 0, \qquad \pi^r &= \mathcal{O}(r^0), \qquad \pi^A &= \mathcal{O}(\frac{1}{r^2}) \end{aligned}$$

- Boundary fields: $A_{(0)A} = A_A|_{r \to \infty}$, $\pi^r_{(0)} = \pi^r|_{r \to \infty}$
- Extended Hamiltonian $\mathcal{H}_E = \mathcal{H}_C + \lambda_u \pi^u + \lambda_A \chi^A + \lambda \partial_i \pi^i$
- Hamilton's equations

$$\dot{A}_{u} = \lambda_{u}, \quad \dot{A}_{r} = \frac{e^{2}\pi^{r}}{r^{2}\sqrt{\gamma}} - \partial_{r}\lambda, \quad \dot{A}_{A} = \frac{e^{2}\tilde{\pi}_{A}}{\sqrt{\gamma}} - \partial_{A}\lambda + \epsilon\lambda_{A}$$

$$\dot{\pi}^{u} = 0, \quad \dot{\pi}^{r} = \frac{\sqrt{\gamma}}{e^{2}}\nabla_{A}\lambda^{A}, \quad \dot{\pi}^{A} = -\frac{\sqrt{\gamma}}{e^{2}}\left(\frac{1}{r^{2}}\nabla_{B}\tilde{F}^{AB} + \partial_{r}\lambda^{A}\right)$$

Fall-off of the multipliers

 $\lambda_u = \mathcal{O}(\frac{1}{r}), \qquad \Lambda_A = \mathcal{O}\left(r^0\right), \qquad \lambda = \underbrace{\mathcal{O}}(\frac{1}{r})_{\text{ind} r} + \operatorname{ind} r = \operatorname{ind} r$

(PUCV)

16 / 30

SUMMARY

| 1 st class constr. | Multipliers | Parameters | Generators | Charges |
|--|--|-----------------------------------|--------------------|---|
| $\pi^{u},\;\chi=\partial_{i}\pi^{i}\ \chi^{\mathcal{A}}_{(0)}$ | $\lambda_{u}, \ \lambda \ \Lambda_{\mathcal{A}}(\phi)$ | $	heta_u, \ 	heta \ \eta_A(\phi)$ | $G[heta] S[\eta]$ | $egin{array}{l} Q[heta] \ Q_{s}[\eta] \end{array}$ |

E

DQC

イロト イヨト イヨト イヨト

SUMMARY

| 1 st class constr. | Multipliers | Parameters | Generators | Charges |
|--|--|-----------------------------------|--------------------|---|
| π^{u} , $\chi = \partial_{i}\pi^{i}$ $\chi^{A}_{(0)}$ | $\lambda_u, \ \lambda \Lambda_A(\phi)$ | $	heta_u, \ 	heta \ \eta_A(\phi)$ | $G[heta] S[\eta]$ | $egin{array}{l} Q[heta] \ Q_s[\eta] \end{array}$ |

- Smeared generators
 - $\begin{array}{ll} G[\theta] &= \int \mathrm{d}^{3}x \left(\theta \, \partial_{i} \pi^{i} + \theta_{u} \pi^{u}\right) & \mathrm{U}(1) \text{ symmetry} \\ S[\eta] &= \int \mathrm{d}^{3}x \, \eta_{A} \chi^{A} & \text{asymptotic symmetry} \end{array}$

<ロト <回ト < 臣ト < 臣ト

SUMMARY

| 1 st class constr. | Multipliers | Parameters | Generators | Charges |
|--|--|-----------------------------------|--------------------|---|
| π^{u} , $\chi=\partial_{i}\pi^{i}$ $\chi^{A}_{(0)}$ | $\lambda_{u}, \ \lambda \ \Lambda_{\mathcal{A}}(\phi)$ | $	heta_u, \ 	heta \ \eta_A(\phi)$ | $G[heta] S[\eta]$ | $egin{array}{l} Q[heta] \ Q_{s}[\eta] \end{array}$ |

Smeared generators

 $\begin{array}{ll} G[\theta] &= \int \mathrm{d}^3 x \left(\theta \, \partial_i \pi^i + \theta_u \pi^u \right) & \mathrm{U}(1) \text{ symmetry} \\ S[\eta] &= \int \mathrm{d}^3 x \, \eta_A \chi^A & \text{asymptotic symmetry} \end{array}$

- New symmetry generator
- $\chi^{\mathcal{A}}$ contains $\chi^{\mathcal{A}}_{(0)}$ (first class) and $\chi^{\mathcal{A}}_{(n)}$ $(n\geq 1)$ (second class)
- The basis $\{r^n|n\in\mathbb{Z}\}$ is not complete, $\chi^{\mathcal{A}} o\chi^{\mathcal{A}}_{(n)}$ not invertible
- We have to ensure that $\chi^{\mathcal{A}}_{(n)}~(n\geq 1)$ do not contribute to δ_{η}

- η_A tends fast to $\eta_{(0)A}$, we can take $\eta_A = \eta_{(0)A}(y)$

• Transformation laws of the fields $\delta_{\theta} = \{ , G[\theta] \}, \delta_{\eta} = \{ , S[\eta] \}$

$$\begin{split} \delta_{\theta} A_{\mu} &= \theta_{u} \, \delta^{u}_{\mu} - \partial_{i} \theta \, \delta^{i}_{\mu} \,, \qquad \delta_{\eta} A_{\mu} &= \epsilon \, \eta_{A} \, \delta^{A}_{\mu} \\ \delta_{\theta} \pi^{\mu} &= 0 \,, \qquad \qquad \delta_{\eta} \pi^{\mu} &= \frac{1}{e^{2}} \, \delta^{\mu}_{r} \sqrt{\gamma} \, \nabla_{A} \eta^{A} \end{split}$$

・ロト ・回 ト ・ ヨト ・

• Transformation laws of the fields $\delta_{\theta} = \{ , G[\theta] \}, \delta_{\eta} = \{ , S[\eta] \}$

$$\begin{aligned} \delta_{\theta} A_{\mu} &= \theta_{u} \, \delta_{\mu}^{u} - \partial_{i} \theta \, \delta_{\mu}^{i} , \qquad \delta_{\eta} A_{\mu} &= \epsilon \, \eta_{A} \, \delta_{\mu}^{A} \\ \delta_{\theta} \pi^{\mu} &= 0 , \qquad \qquad \delta_{\eta} \pi^{\mu} &= \frac{1}{e^{2}} \, \delta_{r}^{\mu} \sqrt{\gamma} \, \nabla_{A} \eta^{A} \end{aligned}$$

igstarrow heta : Standard U(1) gauge transformations $\delta_{ heta} A_{\mu} = -\partial_{\mu} heta$

 $\star \eta^A$: Act only on the boundary fields (Goldstone modes)

$$\delta_{\eta}A_{(0)\mu} = \epsilon \eta_A \delta^A_{\mu}, \quad \delta_{\eta}\pi^r_{(0)} = \frac{1}{e^2}\sqrt{\gamma} \nabla_A \eta^A$$

(PUCV)

• Transformation laws of the fields $\delta_{\theta} = \{ , G[\theta] \}, \delta_{\eta} = \{ , S[\eta] \}$

$$\begin{aligned} \delta_{\theta} A_{\mu} &= \theta_{u} \, \delta_{\mu}^{u} - \partial_{i} \theta \, \delta_{\mu}^{i} \,, \qquad \delta_{\eta} A_{\mu} &= \epsilon \, \eta_{A} \, \delta_{\mu}^{A} \\ \delta_{\theta} \pi^{\mu} &= 0 \,, \qquad \qquad \delta_{\eta} \pi^{\mu} &= \frac{1}{e^{2}} \, \delta_{r}^{\mu} \sqrt{\gamma} \, \nabla_{A} \eta^{A} \end{aligned}$$

igstarrow heta : Standard U(1) gauge transformations $\delta_{ heta} {\sf A}_{\mu} = - \partial_{\mu} heta$

 $\star \eta^A$: Act only on the boundary fields (Goldstone modes)

$$\delta_{\eta}A_{(0)\mu} = \epsilon \eta_{A} \delta^{A}_{\mu}, \quad \delta_{\eta}\pi^{r}_{(0)} = \frac{1}{e^{2}}\sqrt{\gamma} \nabla_{A}\eta^{A}$$

- Transformation law of the multipliers
 - $\delta\lambda_{u} = \dot{\theta}_{u}, \qquad \delta\lambda = \dot{\theta} + \mathcal{O}(\frac{1}{r}), \qquad \delta\Lambda_{A} = \epsilon \,\dot{\eta}_{A} + \mathcal{O}(\frac{1}{r})$
- Fall-off of the local parameters

$$\theta_u = \mathcal{O}(\frac{1}{r}), \qquad \theta = \mathcal{O}(r^0), \qquad \eta^A = \mathcal{O}(r^0)$$

Improper transformations: θ₍₀₎, η^A₍₀₎
 [Benguria, Cordero, Teitelboim 1977]

イロト 不得下 イヨト イヨト 二日

Improved generators and charges

• Improved generators

 $\begin{array}{ll} G_Q[\theta] &= G[\theta] + Q[\theta] & (Q[\theta] = \text{ surface term}) \\ S_Q[\eta] &= S[\eta] + Q_s[\eta] & (Q_s[\theta] = \text{ surface term}) \end{array}$

Differentiability

$$\begin{split} \delta G_Q[\theta] &= \int d^3 x \, \left(\frac{G_Q[\theta]}{\delta A_\mu} \, \delta A_\mu + \frac{G_Q[\theta]}{\delta \pi^\mu} \, \delta \pi^\mu \right) \\ \delta S_Q[\eta] &= \int d^3 x \, \left(\frac{S_Q[\eta]}{\delta A_\mu} \, \delta A_\mu + \frac{S_Q[\eta]}{\delta \pi^\mu} \, \delta \pi^\mu \right) \end{split}$$

3

19 / 30

Improved generators and charges

• Improved generators

 $\begin{array}{ll} G_Q[\theta] &= G[\theta] + Q[\theta] & (Q[\theta] = \text{ surface term}) \\ S_O[\eta] &= S[\eta] + Q_s[\eta] & (Q_s[\theta] = \text{ surface term}) \end{array}$

Differentiability

$$\begin{split} \delta G_Q[\theta] &= \int d^3 x \, \left(\frac{G_Q[\theta]}{\delta A_{\mu}} \, \delta A_{\mu} + \frac{G_Q[\theta]}{\delta \pi^{\mu}} \, \delta \pi^{\mu} \right) \\ \delta S_Q[\eta] &= \int d^3 x \, \left(\frac{S_Q[\eta]}{\delta A_{\mu}} \, \delta A_{\mu} + \frac{S_Q[\eta]}{\delta \pi^{\mu}} \, \delta \pi^{\mu} \right) \end{split}$$

• Charges (field-independent local parameters)

 $Q[\theta] = -\oint d^2 y \,\theta \,\pi^r$ $Q_s[\eta] = \frac{1}{a^2} \oint d^2 y \,\sqrt{\gamma} \,\eta^A A_A$

- The integral is over the two-sphere at the infinity.
 - There is an infinite number of global charges, corresponding to the spherical harmonics mode expansion (or Laurent expansion) of the local parameters.

WHAT'S GOING ON?



Evolution is insensitive to the dynamics that is not prescribed on the initial data surface.

[ALEXANDROV, SPEZIALE 2014]

Presentation of H. González, NordGrav @ ICEN, Iquique, Chile, January 2023

Charge algebra

- Reduced phase space: $G_Q[\theta] = 0 + Q[\theta]$, $S_Q[\eta] = 0 + Q_s[\eta]$
- Algebra can be computed from $\delta_{\theta_2} Q[\theta_1] = \{Q[\theta_1], Q[\theta_2]\}$
- General form of the charge algebra: $\{Q[\theta_1], Q[\theta_2]\} = Q[[\theta_1, \theta_2]] + C[\theta_1, \theta_2]$
- In the electromagnetic theory

 $\{Q[\theta_1], Q[\theta_2]\} = 0$ $\{Q_s[\eta_1], Q_s[\eta_2]\} = 0$ $\{Q[\theta], Q_s[\eta]\} = C[\theta, \eta] \neq 0$

• Central charge $C[\theta, \eta] = \frac{1}{e^2} \oint d^2 y \sqrt{\gamma} \eta^A \partial_A \theta$

Charge algebra

- Reduced phase space: $G_Q[\theta] = 0 + Q[\theta]$, $S_Q[\eta] = 0 + Q_s[\eta]$
- Algebra can be computed from $\delta_{\theta_2} Q[\theta_1] = \{Q[\theta_1], Q[\theta_2]\}$
- General form of the charge algebra: $\{Q[\theta_1], Q[\theta_2]\} = Q[[\theta_1, \theta_2]] + C[\theta_1, \theta_2]$
- In the electromagnetic theory

 $\{Q[\theta_1], Q[\theta_2]\} = 0$ $\{Q_s[\eta_1], Q_s[\eta_2]\} = 0$ $\{Q[\theta], Q_s[\eta]\} = C[\theta, \eta] \neq 0$

- Central charge $C[\theta, \eta] = \frac{1}{e^2} \oint d^2 y \sqrt{\gamma} \eta^A \partial_A \theta$
- Holographic conjugate pairs on S² [Donnay,Puhm, Strominger 2019] $\{Q[\theta], Q_s[\eta]\} = C[\theta, \eta] \leftrightarrow \{q, p\} = 1$ $Q[\theta]$ - conformally soft photon mode $Q_s[\eta]$ - Goldstone current

Algebra in modes

• Laurent series

$$\psi(z, \bar{z}) = \sum_{n=-\infty}^{+\infty} \sum_{m=-\infty}^{+\infty} \frac{\psi_{nm}}{z^{n+h} \bar{z}^{m+\bar{h}}}$$

- The powers $(h, ar{h})$ are related to the spin of the tensor ψ
- Scalars $\pi^r:(0,0)$
- Vectors: A_z : (1, 0), $A_{\bar{z}}$: (0, 1)

(PUCV)

Algebra in modes

• Laurent series

$$\psi(z, \bar{z}) = \sum_{n=-\infty}^{+\infty} \sum_{m=-\infty}^{+\infty} \frac{\psi_{nm}}{z^{n+h} \bar{z}^{m+\bar{h}}}$$

- The powers $(h, ar{h})$ are related to the spin of the tensor ψ
- Scalars $\pi^r:(0,0)$
- Vectors: A_z : (1, 0), $A_{\bar{z}}$: (0, 1)
- Charges

 $Q[\theta] = \sum_{n,m} \theta_{nm} G_{nm} \in \mathbb{R}$ $Q_s[\eta] = \sum_{n,m} (\eta_{nm} \bar{S}_{nm} + \bar{\eta}_{nm} S_{nm}) \in \mathbb{R}$

Generators

$$G_{nm} = 4\pi^2 \pi_{1-n,1-m}$$

$$S_{nm} = -\frac{4\pi^2}{e^2} A_{-n,-m}$$

$$\bar{S}_{nm} = -\frac{4\pi^2}{e^2} \bar{A}_{-n,-m}$$

< D > < P > < P > < P >

• Algebra (non vanishing brackets only)

$$\{G_{nm}, S_{kl}\} = \kappa n \,\delta_{n+k,0} \delta_{m+l,0}$$

$$\{G_{nm}, \bar{S}_{kl}\} = \kappa m \,\delta_{n+k,0} \delta_{m+l,0}$$

• Level of the algebra: $\kappa = \frac{4\pi^2}{e^2}$

• Algebra (non vanishing brackets only)

 $\{G_{nm}, S_{kl}\} = \kappa n \,\delta_{n+k,0} \delta_{m+l,0}$ $\{G_{nm}, \bar{S}_{kl}\} = \kappa m \,\delta_{n+k,0} \delta_{m+l,0}$

- Level of the algebra: $\kappa = \frac{4\pi^2}{e^2}$
- Change of the basis: $(G_{nm}, S_{nm}, \overline{S}_{nm}) \rightarrow (R_{nm}, J_{nm}, \overline{J}_{nm})$
- Generalization of Kac-Moody algebra

$$\{J_{nm}, J_{kl}\} = \kappa (n-m) \, \delta_{n+k,0} \delta_{m+l,0}$$

$$\{\bar{J}_{nm}, \bar{J}_{kl}\} = -\kappa (n-m) \, \delta_{n+k,0} \delta_{m+l,0}$$

$$\{R_{nm}, J_{kl}\} = \kappa n \, \delta_{n+k,0} \delta_{m+l,0}$$

$$\{R_{nm}, \bar{J}_{kl}\} = \kappa (n+m) \, \delta_{n+k,0} \delta_{m+l,0}$$

Abelian Kac-Moody subalgebras

• We obtain six Abelian KM algebras $\{j_n, j_m\} = \kappa n \, \delta_{n+m,0}$

Currents j_n Levels J_{n0}, J_{0n} $\kappa, -\kappa$ $\bar{J}_{n0}, \bar{J}_{0n}$ $-\kappa, \kappa$ R_{n0}, R_{0n} κ, κ

- Non vanishing mixed brackets: $\{R_{n0}, J_{m0}\}, \{R_{0n}, \overline{J}_{0m}\} \neq 0$
- Each KM algebra is naturally generated by current proportional to a holomorphic or anti-holomorphic functions.

イロト イポト イヨト イ

Abelian Kac-Moody subalgebras

• We obtain six Abelian KM algebras $\{j_n, j_m\} = \kappa n \, \delta_{n+m,0}$

Currents j_n Levels J_{n0}, J_{0n} $\kappa, -\kappa$ $\bar{J}_{n0}, \bar{J}_{0n}$ $-\kappa, \kappa$ R_{n0}, R_{0n} κ, κ

- Non vanishing mixed brackets: $\{R_{n0}, J_{m0}\}, \{R_{0n}, \overline{J}_{0m}\} \neq 0$
- Each KM algebra is naturally generated by current proportional to a holomorphic or anti-holomorphic functions.
- $\{J_{00}, \overline{J}_{00}, R_{00}\}$ span the global Abelian algebra

イロト イポト イヨト イ

Beyond U(1) – conformal symmetry

- Conformal plane a realization of conformal symmetry described by Virasoro algebra
- Witt algebra obtained from KM algebra using the Sugawara construction [Sugawara 1967]

イロト イポト イヨト イヨト

Beyond $U(1)\mbox{--}$ conformal symmetry

- Conformal plane a realization of conformal symmetry described by Virasoro algebra
- Witt algebra obtained from KM algebra using the Sugawara construction [Sugawara 1967]
- We have several Witt algebras, not all of them independent
- Six Witt generators $L_n = \frac{1}{2\kappa} \sum_k j_k j_{n-k}$
- Six Witt algebras $\{L_n, L_m\} = (n-m) L_{n+m}$
- Quantization introduces a central extension.

イロト イポト イヨト イヨト

Beyond $U(1)\mbox{--}$ conformal symmetry

- Conformal plane a realization of conformal symmetry described by Virasoro algebra
- Witt algebra obtained from KM algebra using the Sugawara construction [Sugawara 1967]
- We have several Witt algebras, not all of them independent
- Six Witt generators $L_n = \frac{1}{2\kappa} \sum_k j_k j_{n-k}$
- Six Witt algebras $\{L_n, L_m\} = (n-m) L_{n+m}$
- Quantization introduces a central extension.
- In progress: Relation to the global 4D Poincaré generators.

イロト 不得下 イヨト イヨト

Extension to Yang-Mills theory

• Zero mode

 $\partial_r V_A = -[A_r, V_A] \quad \Rightarrow \quad V_A(x) = U^{-1} V_{(0)A}(y) U,$

with $U = \exp\left(-\int_r^\infty \mathrm{d} r \, A_r\right)$ and the bdy. condition $V_A|_{r\to\infty} = V_{(0)A}(y)$

1

Extension to Yang-Mills theory

• Zero mode

 $\partial_r V_A = -[A_r, V_A] \quad \Rightarrow \quad V_A(x) = U^{-1} V_{(0)A}(y) U,$

with $U = \exp\left(-\int_r^\infty \mathrm{d} r \, A_r\right)$ and the bdy. condition $V_A|_{r\to\infty} = V_{(0)A}(y)$

Charges

 $Q[\theta] = -\oint \mathrm{d}^2 y \, heta^a \pi^r_a$, $Q_s[\eta] = rac{1}{g^2} \oint \mathrm{d}^2 y \, \sqrt{\gamma} \, \eta^a_A A^A_a$

Local transformations

 $\begin{aligned} \delta_{\theta,\eta} A_u^a &= \theta_u^a \qquad \delta_{\theta,\eta} A_r^a = -D_r \theta^a \\ \delta_{\theta,\eta} A_A^a &= -D_A \theta^a + \epsilon \, \eta_A^a \end{aligned}$

▲ロト ▲圖ト ▲ヨト ▲ヨト ニヨー のへで

Extension to Yang-Mills theory

Zero mode

 $\partial_r V_A = -[A_r, V_A] \quad \Rightarrow \quad V_A(x) = U^{-1} V_{(0)A}(y) U,$

with $U = \exp\left(-\int_r^\infty \mathrm{d} r \, A_r\right)$ and the bdy. condition $V_A|_{r\to\infty} = V_{(0)A}(y)$

Charges

 $Q[\theta] = -\oint \mathrm{d}^2 y \, \theta^a \pi^r_a$, $Q_s[\eta] = rac{1}{g^2} \oint \mathrm{d}^2 y \, \sqrt{\gamma} \, \eta^a_A A^A_a$

• Local transformations

$$\begin{split} \delta_{\theta,\eta} A_u^a &= \theta_u^a \qquad \delta_{\theta,\eta} A_r^a = -D_r \theta^a \\ \delta_{\theta,\eta} A_A^a &= -D_A \theta^a + \epsilon \, \eta_A^a \end{split}$$

• Non-Abelian charge algebra

 $\{Q[\theta_1], Q[\theta_2]\} = Q[[\theta_1, \theta_2]]$ $\{Q[\theta], Q_s[\eta]\} = Q_s[[\theta, \eta]] + \frac{1}{g^2} \oint d^2 y \sqrt{\gamma} \eta_a^A \partial_A \theta^a$ $\{Q_s[\eta_1], Q_s[\eta_2]\} = 0 \longrightarrow Q_s \text{ is Abelian}$

Yang-Mills theory

$$I[A] = -\frac{1}{4g^2} \int \mathrm{d}^4 x \sqrt{\mathfrak{g}} \, F^{\mu\nu}_a F^a_{\mu\nu}; \quad F^a_{\mu\nu} = \partial_\mu A^a_\nu - \partial_\nu A^a_\mu + f^a_{\ bc} A^b_\mu A^c_\nu$$

• Constraints

$$\pi_a^u \approx 0$$
, $\chi_a^A \equiv \epsilon \, \pi_a^A - \frac{1}{g^2} \sqrt{\gamma} \gamma^{AB} F_{rB}^a \approx 0$, $\chi_a \equiv D_i \pi_a^i \approx 0$

イロト イヨト イヨト イヨト

Yang-Mills theory

$$I[A] = -\frac{1}{4g^2} \int \mathrm{d}^4 x \sqrt{\mathfrak{g}} \, F^{\mu\nu}_a F^a_{\mu\nu}; \quad F^a_{\mu\nu} = \partial_\mu A^a_\nu - \partial_\nu A^a_\mu + f^a_{\ bc} A^b_\mu A^c_\nu$$

• Constraints

$$\pi^{\rm u}_a\approx 0\,,\quad \chi^{\rm A}_a\equiv \epsilon\,\pi^{\rm A}_a-\tfrac{1}{g^2}\,\sqrt{\gamma}\gamma^{AB}F^a_{rB}\approx 0,\quad \chi_a\equiv D_i\pi^i_a\approx 0$$

• Constraint algebra

$$\begin{cases} \chi_a, \chi'_b \} &= f_{ab}^{\ c} \chi_c \, \delta^{(3)} \\ \\ \left\{ \chi_a, \chi'_b^A \right\} &= f_{ab}^{\ c} \, \chi_c^A \, \delta^{(3)} \\ \\ \left\{ \chi_a^A, \chi'_b^B \right\} &= \Omega^{AB}_{ab}(x, x')$$

• Non-Abelian symplectic matrix

$$\Omega_{ab}^{AB}(\mathbf{x},\mathbf{x}') = -\frac{2\epsilon}{g^2} \sqrt{\gamma} \gamma^{AB} \left(g_{ab} \partial_r + f_{abc} A_r^c \right) \delta^{(3)}$$

25 / 30

(PUCV)

Mode algebra

$$\begin{cases} G_{nm}^{a}, G_{kl}^{b} \\ G_{nm}^{a}, S_{kl}^{b} \end{cases} = f_{c}^{ab} G_{n+k,m+l}^{c}$$
$$\begin{cases} G_{nm}^{a}, S_{kl}^{b} \\ G_{nm}^{a}, \bar{S}_{kl}^{b} \end{cases} = f_{c}^{ab} \bar{S}_{n+k,m+l}^{c} + \kappa n g^{ab} \delta_{n+k,0} \delta_{m+l,0}$$
$$\begin{cases} G_{nm}^{a}, \bar{S}_{kl}^{b} \\ G_{nm}^{a}, \bar{S}_{kl}^{b} \end{cases} = f_{c}^{ab} \bar{S}_{n+k,m+l}^{c} + \kappa n g^{ab} \delta_{n+k,0} \delta_{m+l,0}$$

- Level $\kappa = \frac{4\pi^2}{g^2}$
- One can apply the Sugawara method again...

Mode algebra

$$\begin{cases} G_{nm}^{a}, G_{kl}^{b} \\ = f_{c}^{ab} G_{n+k,m+l}^{c} \\ \begin{cases} G_{nm}^{a}, S_{kl}^{b} \\ \end{cases} = f_{c}^{ab} S_{n+k,m+l}^{c} + \kappa n g^{ab} \delta_{n+k,0} \delta_{m+l,0} \\ \begin{cases} G_{nm}^{a}, \bar{S}_{kl}^{b} \\ \end{cases} = f_{c}^{ab} \bar{S}_{n+k,m+l}^{c} + \kappa m g^{ab} \delta_{n+k,0} \delta_{m+l,0} \end{cases}$$

- Level $\kappa = \frac{4\pi^2}{g^2}$
- One can apply the Sugawara method again...
- We conclude that the symmetries at the asymptotic null boundary, described by KM algebras and Virasoro algebras, are general features of 4D gauge theories

Asymptotic conditions

- Invariance of boundary conditions under Poincaré transformatons is not straighforward
- Hamiltonian treatment at spatial infinity needs additional **parity conditions** to ensure invariance under boosts.
- Electromagnetism [Henneaux, Troessaert 2018] ;
- Yang-Mills [Tanzi, Giulini 2020]
- Null-slices foliated standard b.c. in electromagnetism are invariant under Poincaré group. [Bunster, Gomberoff, Pérez 2018]
- We showed the Poincaré invariance in the non-Abelian case.

Poincaré transfromations

- We found several Kac-Moody algebras, but not all of them are related to the global Poincaré symmetry in 4D spacetime.
- We constructed a generator of 4D Poincaré transformations and its action at the light front, by writing the YM stress tensor in the canonical form,

$$T^{\mu}_{\nu} = \frac{1}{g^2} \left(F^{\mu\alpha}_{a} F^{a}_{\nu\alpha} - \frac{1}{4} \,\delta^{\mu}_{\nu} \,F^{\alpha\beta}_{a} F^{a}_{\alpha\beta} \right)$$

• We are working on showing its relation with the Virasoro generators.

Image: A match a ma
Poincaré transfromations

- We found several Kac-Moody algebras, but not all of them are related to the global Poincaré symmetry in 4D spacetime.
- We constructed a generator of 4D Poincaré transformations and its action at the light front, by writing the YM stress tensor in the canonical form,

$$T^{\mu}_{\nu} = \frac{1}{g^2} \left(F^{\mu\alpha}_{a} F^{a}_{\nu\alpha} - \frac{1}{4} \, \delta^{\mu}_{\nu} \, F^{\alpha\beta}_{a} F^{a}_{\alpha\beta} \right)$$

• We are working on showing its relation with the Virasoro generators.

To be done

- Hamiltonian treatment of the gravitatonal action using the null foliation.
- Description of a holographic theory.
- Addition of the θ -term in the action (Pontryagin topological invariant with the couplig θ), which will change the central charges.

29 / 30

(PUCV)

THANK YOU!

30 / 30

Acknowledgments

THANK YOU!



Acknowledgments

FONDECYT Grant N° 1190533 Anillo Grant ANID/ACT210100 Holography and its applications to high energy physics, quantum gravity and condensed matter systems

FONDECYT Grant No. 1190533 Black holes and asymptotic symmetries

30 / 30

イロト イポト イヨト イヨト

(PUCV)